

Ph136 Solution 6

$$\begin{aligned}
 6.6 \quad (a) \quad \rho(r) &= \langle \rho(\vec{x}) \rho(\vec{x} + \vec{r}) \rangle = \lim_{V \rightarrow \infty} \frac{1}{V} \int_V dx \delta x \rho(x+r) \\
 &= \lim_{V \rightarrow \infty} \frac{1}{V} \frac{1}{(2\pi)^3} \int d\vec{k} e^{i\vec{k} \cdot \vec{r}} \rho_V(\vec{k}) \rho_V(-\vec{k}) \\
 &= \int \frac{d\vec{k}}{(2\pi)^3} e^{-i\vec{k} \cdot \vec{r}} P(k)
 \end{aligned}$$

The universe is isotropic $\Rightarrow \rho(\vec{x}) = \rho(|\vec{x}|) \Rightarrow P(\vec{k}) = P(k)$

Using the fact $\int_0^\pi d\theta \sin\theta \exp(-ikr \cos\theta) = 2 \text{sinc}(kr)$

$$\Rightarrow \rho(r) = \int_0^\infty \frac{dk}{(2\pi)^2} k^2 \text{sinc}(kr) P(k)$$

(b) Trivial

$$(c) \quad \rho_R(\vec{x}) = \frac{3}{4\pi R^3} \int_{r < R} dr \delta(x+r) = \int_V dr \delta(x+r) k(r)$$

$$\text{where } k(r) = \begin{cases} \frac{3}{4\pi R^3} & \text{if } r < R \\ 0 & \text{elsewhere} \end{cases}$$

$$\text{So } \sigma^2 = \langle \rho_R^2(x) \rangle = \lim_{V \rightarrow \infty} \int \frac{1}{V} \frac{d\vec{k}}{(2\pi)^3} |\rho_R(\vec{k})|^2$$

By convolution theorem, $|\rho_R(k)|^2 = |\rho_V(k)|^2 |k(k)|^2$

Also using spherical coordinates, $\tilde{k}(k) = W(kR)$, where $W(x) = \frac{3(\text{sinc} x - x \cos x)}{x^2}$

$$\Rightarrow \sigma^2 = \int_0^\infty \frac{dk}{(2\pi)^2} k^2 P(k) W^2(kR)$$

$$6.10 \quad (a) \quad N(t) = \int K(t'-t) Y(t') dt' = \int K(t'') Y(t''+t) dt'' \\ = \int K(t'') Y(t'') dt''$$

where we made a change of variable $t'' = t' - t$ and noted that Y is a stationary random process. Thus, the statistical property is independent of time. Also $\overline{N(t)} = \langle N(t) \rangle = \int K \bar{y} dt = 0$

$$\Rightarrow \overline{N^2(t)} = \sigma_N^2 = \int_0^\infty S_N(f) df = \int_0^\infty |K(f)|^2 S_Y(f) df$$

(b) Note $K(-f) = K^*(f)$, $S(-f) = S^*(f)$

$$S = \int_0^\infty K(f) S^*(f) df + c.c$$

$$\Rightarrow \frac{S}{\langle N^2 \rangle^{1/2}} = \frac{\int_0^\infty K(f) S^*(f) df + c.c}{\left(\int_0^\infty |K(f)|^2 S_Y(f) df \right)^{1/2}}, \text{ Vary } K(f) \rightarrow K(f) + \delta K(f)$$

$$\Rightarrow \delta \left(\frac{S}{\langle N^2 \rangle^{1/2}} \right) = \frac{S}{\langle N^2 \rangle^{1/2}} \left[\int_0^\infty df \delta K \left(\frac{S(f)}{S} - \frac{K(f) S_Y(f)}{2 \langle N^2 \rangle} \right) + c.c \right]$$

$$\Rightarrow K(f) = \text{const} \times \frac{S(f)}{S_Y(f)}$$

$K(f)$ is maximum is obvious, which can be checked by doing 2nd derivative.

$$\begin{aligned}
6.10 \text{ (c)} \quad \frac{S^2}{\langle N^2 \rangle} &= \frac{C \int kcf) S^*cf) df + c.c.)^2 df}{\int |kcf)|^2 Sycf) df} \\
&= \frac{(\int_0^\infty (\frac{|Scf)|^2}{Sycf) + c.c.) df)^2}{\int_0^\infty \frac{S^2cf)}{Sycf) df} \\
&= 4 \int_0^\infty \frac{|Scf)|^2}{Sycf) df}
\end{aligned}$$

7.2 (a) Taylor expanding $\Omega(k)$ to 2nd order and noticing

$$v_g = \frac{d\Omega}{dk} \Big|_{k_0} \Rightarrow \Omega = \omega_0 + v_g k + (dv_g/dk) k^2/2$$

(b) Trivial

(c) Note

$$\int_{-\infty}^{\infty} dk \exp(- (a^2 k^2 + i b k + i c k^2)) = \left(\frac{\pi}{a^2 + i c} \right)^{1/2} \exp\left(- \frac{b^2 (a^2 - i c)}{4(a^2 + i c)} \right)$$

$$\text{Now } a^2 = \frac{1}{2\Delta k^2}, \quad b = -(x - x_0 - v_g t), \quad c = \frac{1}{2} \frac{dv_g}{dk} t$$

$$\Rightarrow |A| \propto \exp\left[- \frac{b^2 a^2}{4(a^2 + i c)} \right] = \exp\left[- \frac{(x - x_0 - v_g t)^2}{2L^2} \right]$$

$$\text{with } L = \frac{1}{\Delta k} \sqrt{1 + \left(\frac{dv_g}{dk} 2k^2 t \right)^2}$$

7.2 (d) At $t=0$, $L = \frac{1}{\Delta k}$. Recalling that $L \sim \Delta x$

$\Rightarrow \Delta x \Delta k \sim 1$ (uncertainty principle).

(e) $L(t=0) = \frac{1}{\Delta k}$, Solving $L(t) = 2L(0)$

$\Rightarrow t = \frac{\sqrt{3}}{\frac{dv_s}{dk} (\Delta k)^2}$. Now let's consider a wave packet travelling

from Hawaii to California. Spreading less than 2 $\Rightarrow \left(\frac{dv_s}{dk}\right) \frac{1}{\Delta x^2} t < \sqrt{3}$

Also for waves on the surface of deep water: $\omega = \sqrt{gk}$

And $t = \frac{D}{v_s}$ where D being the distance from Hawaii \rightarrow California

$D \sim 3 \times 10^3 \text{ km}$, $k_0 = \frac{2\pi}{\lambda_0}$ and $\lambda_0 \sim 100 \text{ m}$

$$\Rightarrow \Delta x > \sqrt{\frac{D \lambda_0}{4\pi\sqrt{3}}} \sim 4 \text{ km} = 40 \lambda_0$$